

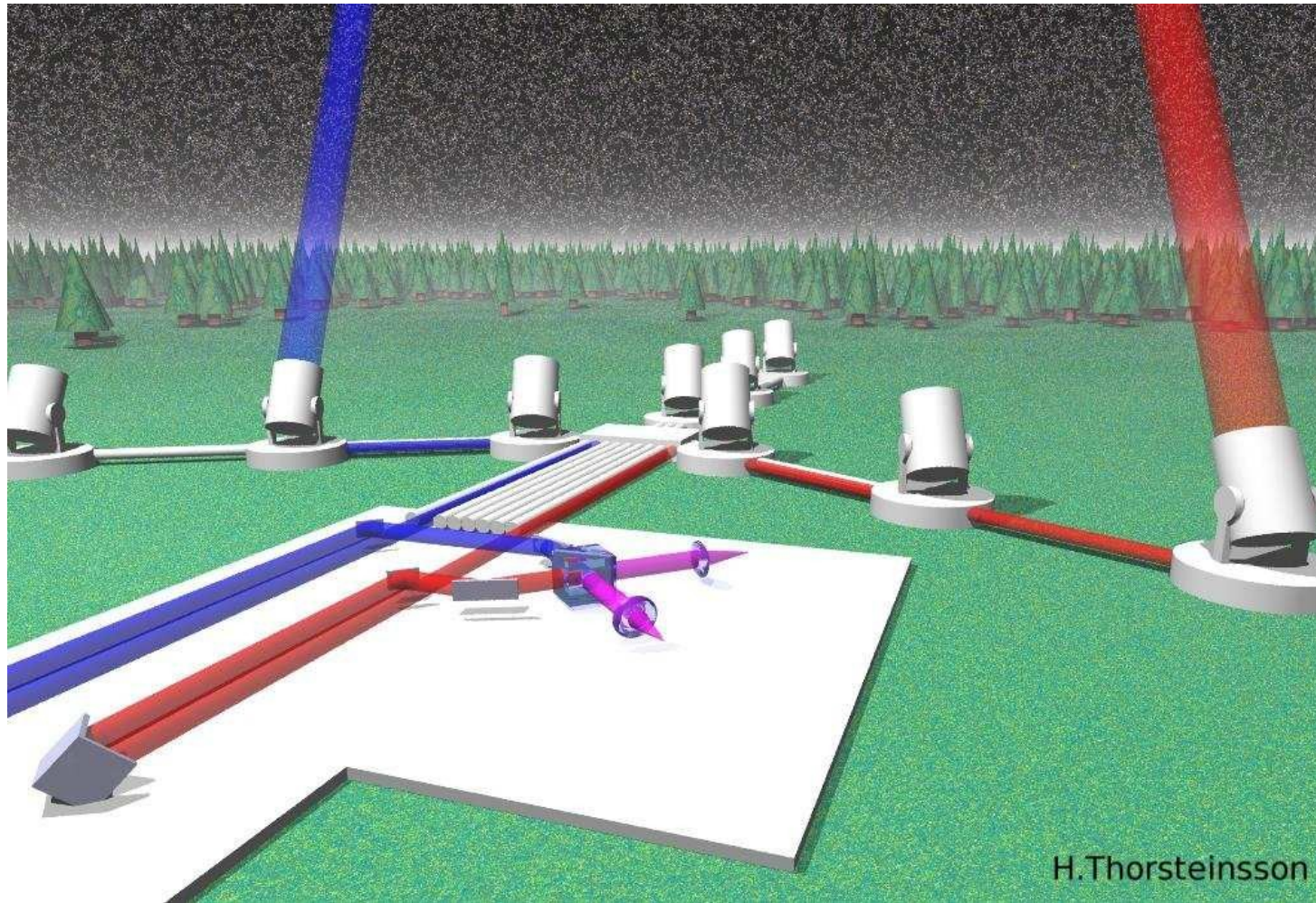


UNIVERSITY OF
CAMBRIDGE

What kinds of interferometric science will be
feasible from the ground?

David Buscher
Cavendish Laboratory
University of Cambridge

We would like to develop science cases for
a new interferometer



A purely “science-driven” approach can develop science cases which are infeasible



Understanding the technical limitations will serve to focus the science



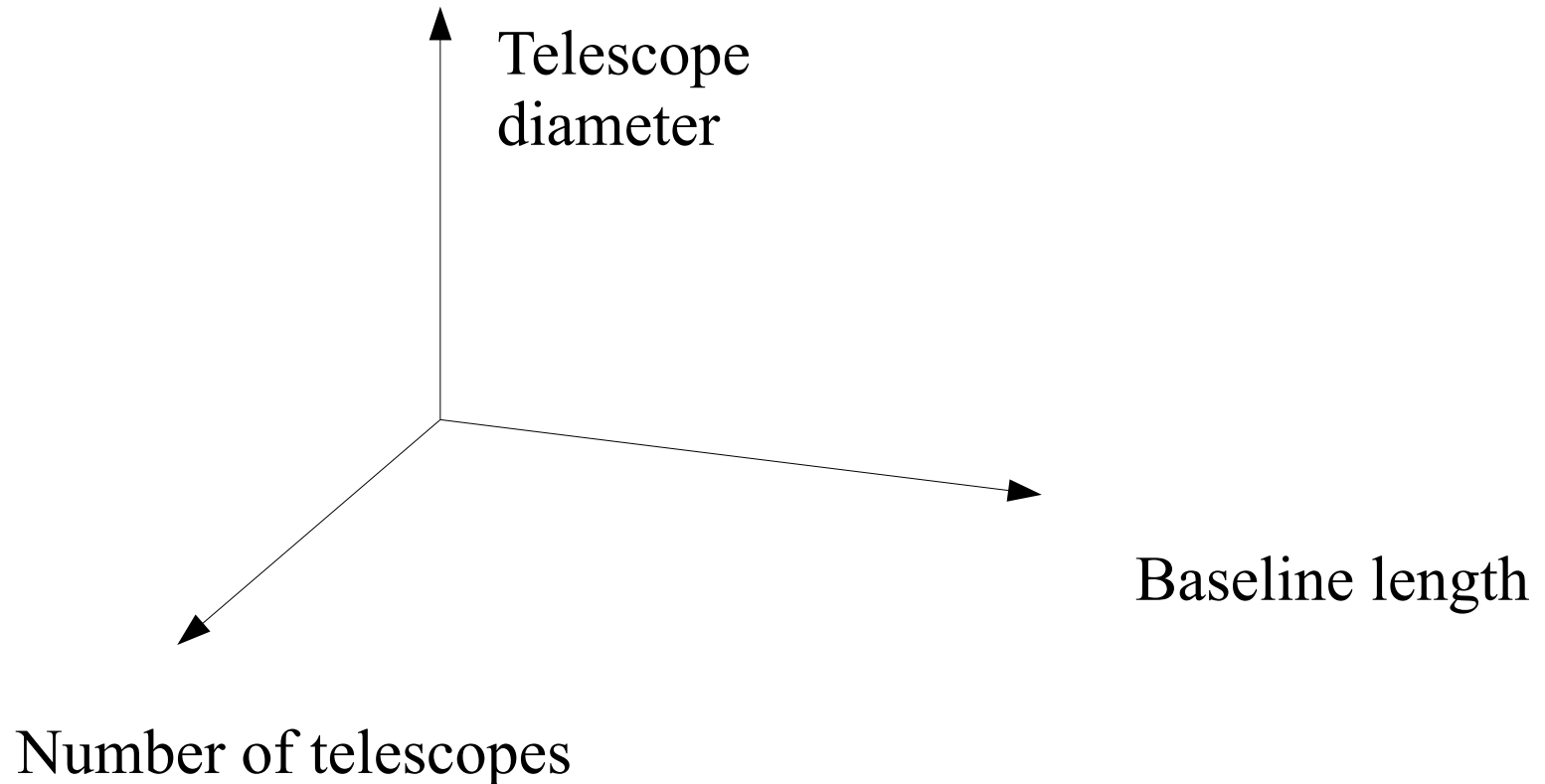
By mapping out the “feasibility landscape”
we can plan a route to the future



A new interferometer must have significant new capability to do new science

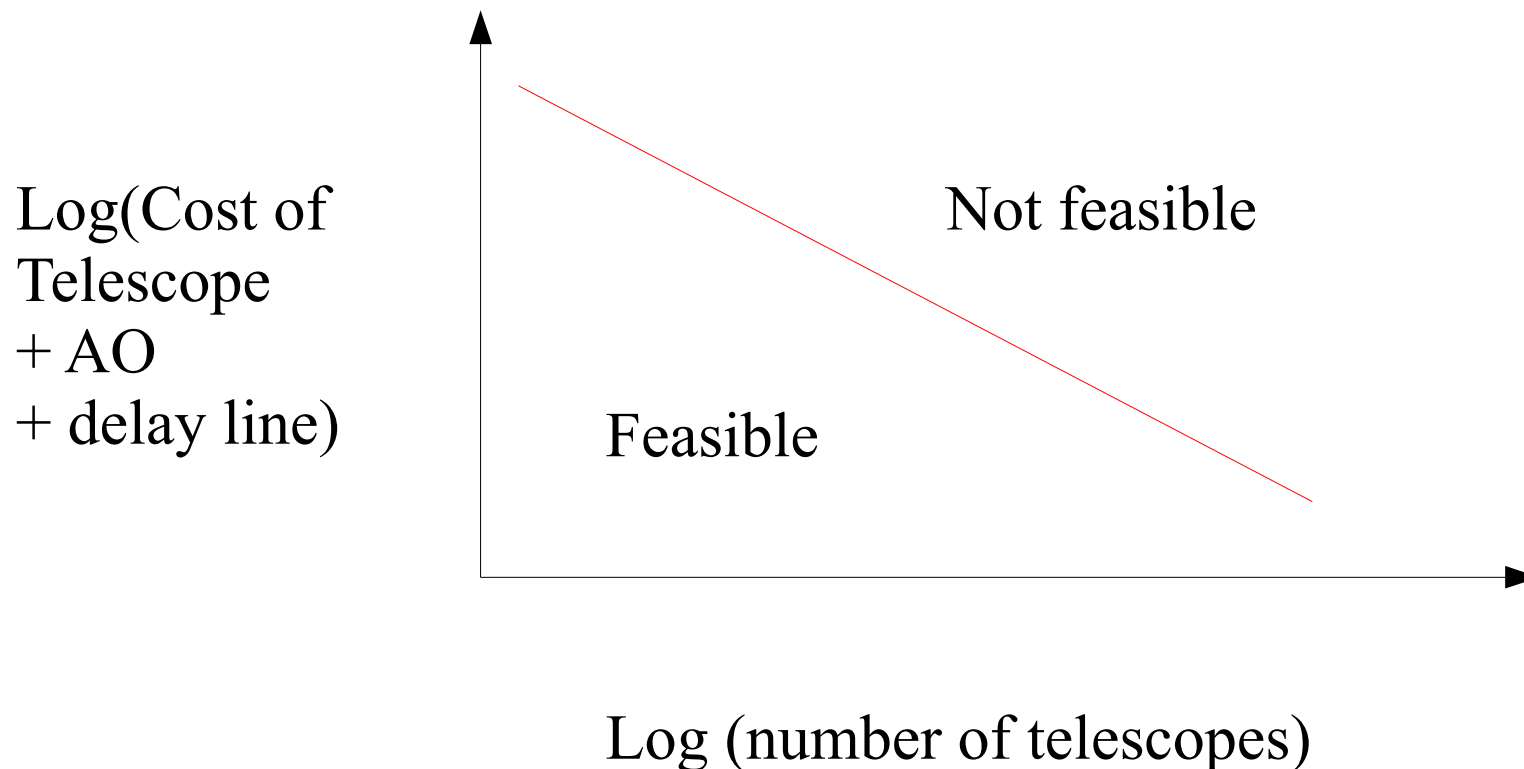
- Spatial resolution
 - Baselines of many km to resolve X-ray binaries
- Spectral resolution
 - $R \sim 30000$ velocity-resolved stellar winds
- Dynamic range
 - Better than 10^4 needed for hot Jupiters
- Limiting magnitude
 - More targets
- Image complexity
 - Complex physics – turbulence, clumping

We can represent possible interferometers as points in a multi-dimensional space

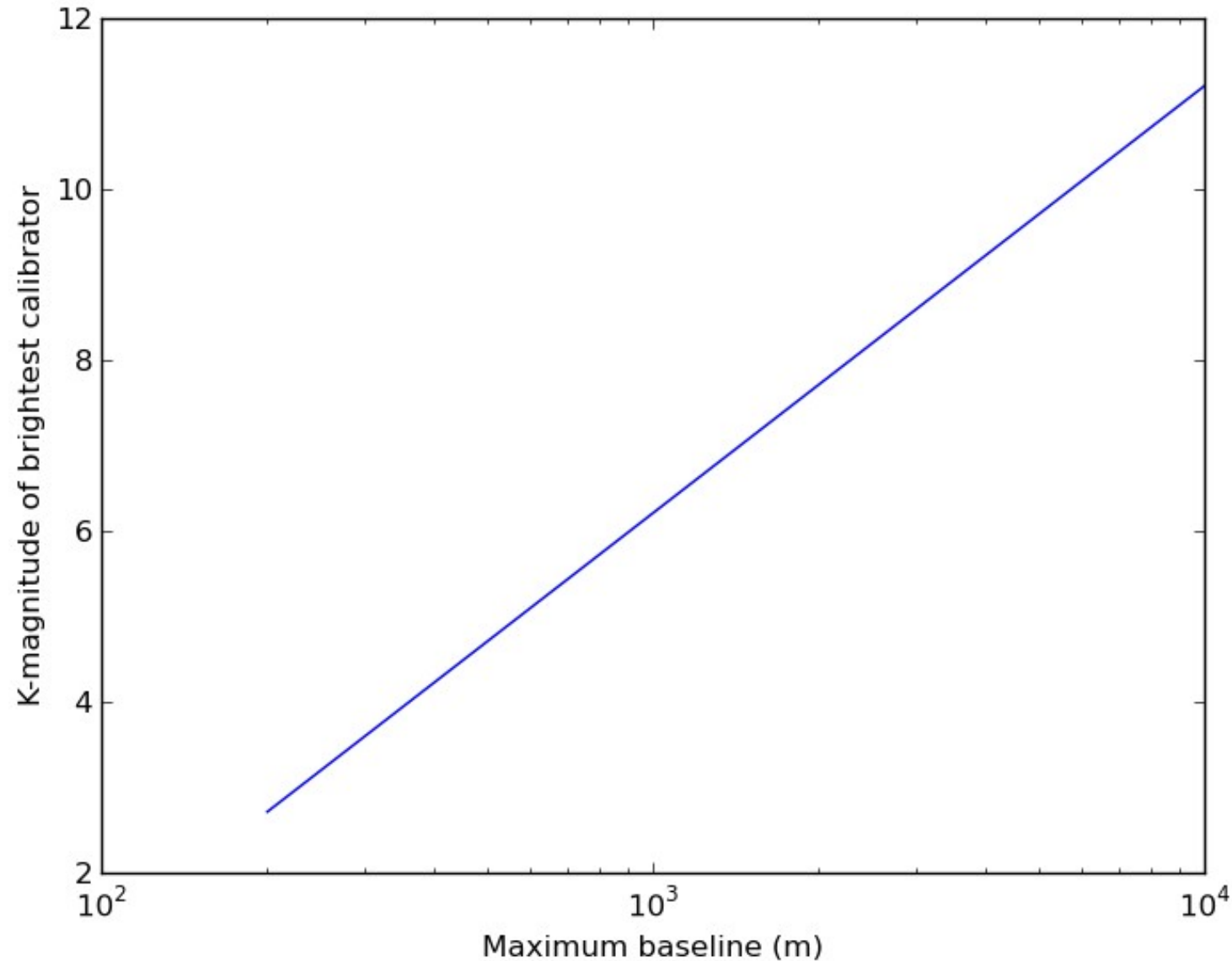


Constraints on funding and timescales mean that we cannot do everything

- Funding < \$1 billion?
- Operational by 2030?

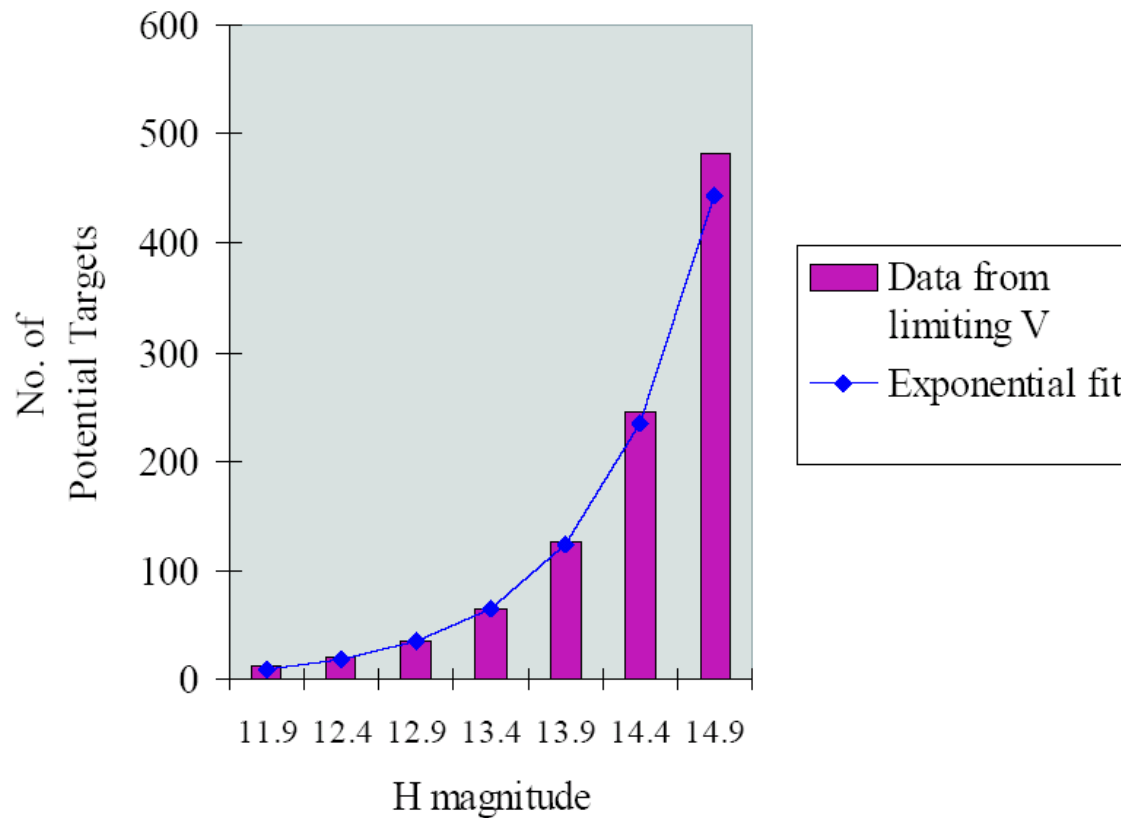


Fundamental factors also serve to make our science parameter choices non-orthogonal



Increasing the limiting magnitude would allow access to many science targets

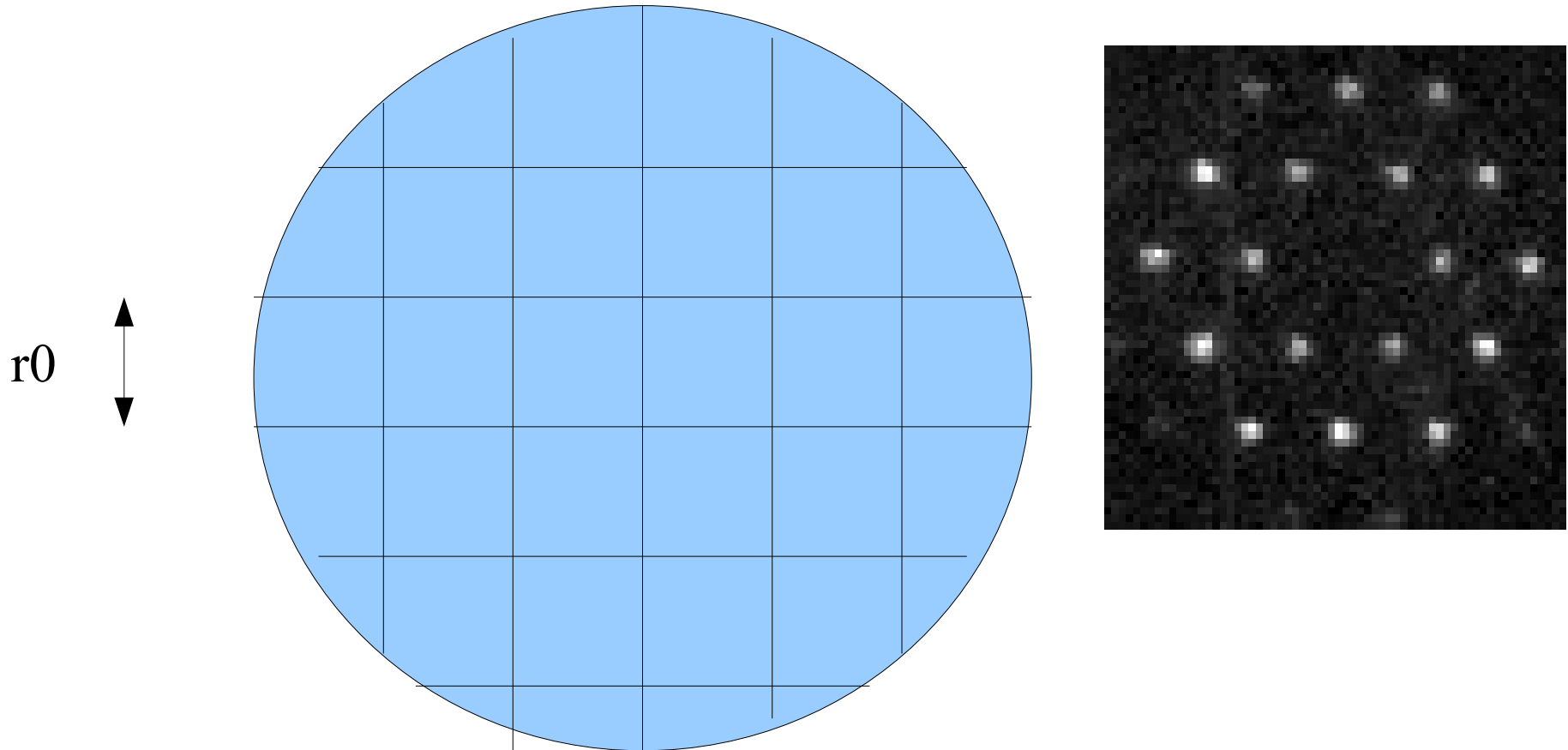
Potential Targets vs H magnitude



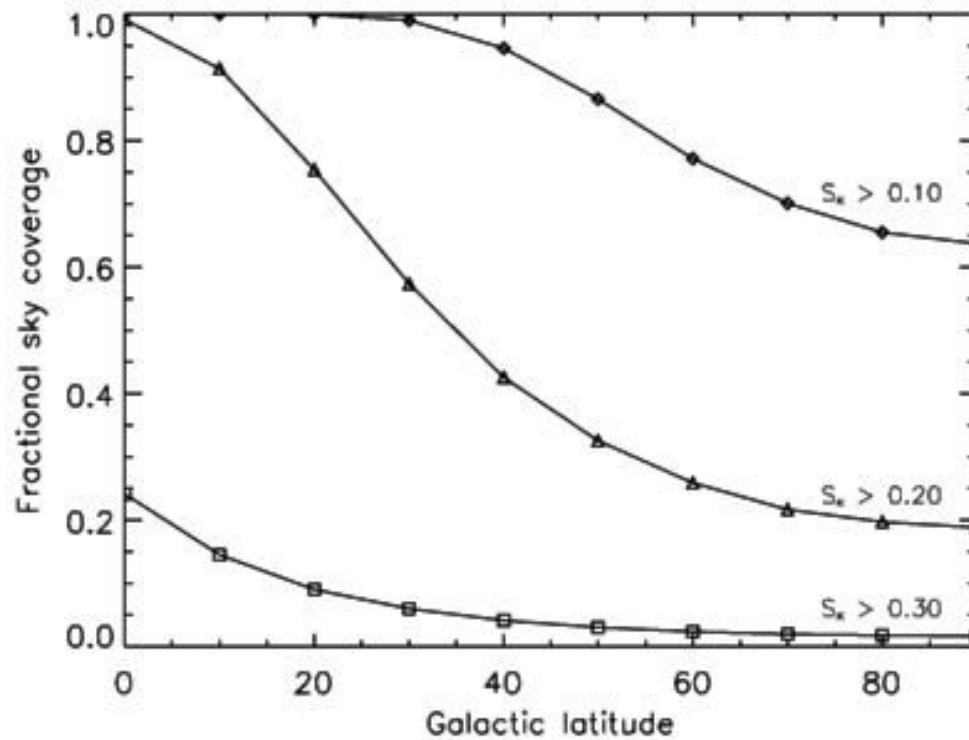
Perhaps throw money at large telescopes
and lasers?



In practice we are still limited by the
photons per r_0 patch per t_0

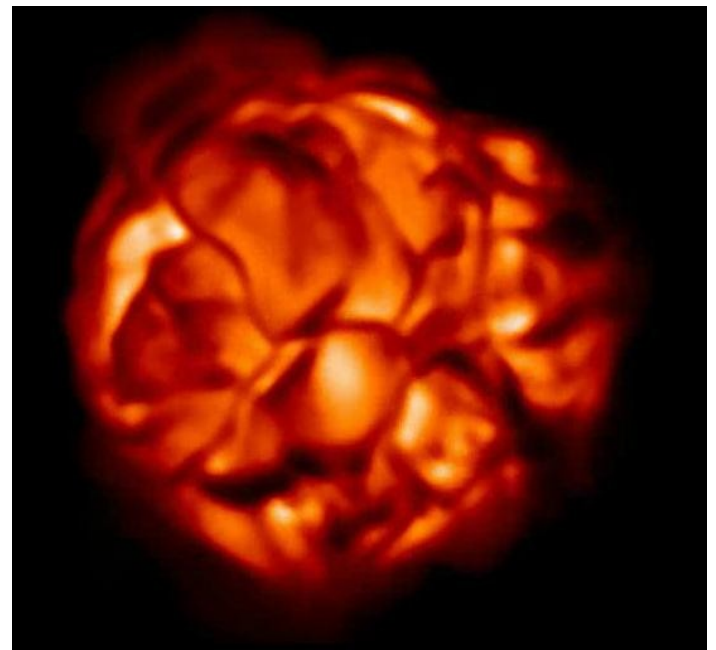


In reality there is no “magic bullet”



Increasing the number of telescopes offers a range of science advantages

- Number of baselines \sim
Number of telescopes²
- Number of “filled pixels” in image \sim Number of baselines
- Dynamic range \sim
SNR * $\sqrt{\text{Number of baselines}}$
- Baseline “bootstrapping” on extended objects



A straw-man proposal

- Increasing the limiting magnitude is not cost-effective versus buying more telescopes
- Can use self-referenced AO and modest sized telescopes to increase the magnitude at which we can “freeze the fringes” to perhaps $K \sim 13-14$
- With good fringe tracking can do higher dynamic range, better spectral resolution.

Need to iterate between technology and science

